

# ELL212 - Tutorial 4 Solutions, Sem II 2015-16

## Problem 1

a) The charge density  $\rho(r)$  is given by:

$$\rho(r) = e(p(r) - n(r)) = en_0 \left( \exp\left(-\frac{e\phi(r)}{k_B T}\right) - \exp\left(\frac{e\phi(r)}{k_B T}\right) \right) \quad (1)$$

For small  $Q$ , the potential  $\phi(r)$  set-up in the plasma is also small, and thus the charge density can be approximated by a linear expression in  $\phi(r)$ :

$$\rho(r) = -\frac{2e^2 n_0}{k_B T} \phi(r) \quad (2)$$

Substituting into the Poisson's equation:

$$\nabla^2 \phi(r) = \frac{2e^2 n_0}{\epsilon_0 k_B T} \phi(r) \quad (3)$$

Using the spherical coordinate system (wherein for a  $r$  dependent potential,  $\nabla^2 \phi(r) \equiv \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 \frac{\partial}{\partial r} (r^2 \phi(r))) = \frac{1}{r} \frac{\partial^2}{\partial r^2} (r \phi(r))$ ), we obtain: In spherical coordinates, use Laplacian expression,

$$\boxed{\frac{1}{r} \frac{\partial^2}{\partial r^2} (r \phi(r)) = \frac{2e^2 n_0}{\epsilon_0 k_B T} \phi(r)} \quad (4)$$

b) The solution to Eq. 4 can be written as

$$\phi(r) = A_1 \frac{\exp(r/\lambda_D)}{r} + A_2 \frac{\exp(-r/\lambda_D)}{r} \quad (5)$$

where  $\lambda_D = \sqrt{\epsilon_0 k_B T / 2n_0 e^2}$ . Clearly  $A_1 = 0$  since the potential (and hence the electric field) cannot go to  $\infty$  as  $r \rightarrow \infty$ . To compute  $A_2$ , we make use of the fact that there is a point charge  $Q$  sitting at the origin. The electric field at a distance  $r$  from the origin:

$$\mathbf{E}(r) = -\hat{r} \frac{\partial \phi(r)}{\partial r} = \hat{r} \frac{A_2}{r^2} \left( 1 + \frac{r}{\lambda_D} \right) \exp(-r/\lambda_D) \quad (6)$$

Consider a Gaussian surface of a very small radius  $R$  centered at the origin, then the flux of  $\mathbf{E}$  through this surface is given by:

$$\Phi = \int \mathbf{E} \cdot d\mathbf{S} = 4\pi A_2 \left( 1 + \frac{R}{\lambda_D} \right) \exp(-R/\lambda_D) \quad (7)$$

The charge enclosed within this surface is given by

$$Q_{\text{enc}} = Q + 4\pi \int_0^R \rho(r) r^2 dr = Q - \frac{8\pi e^2 n_0}{k_B T} \int_0^R \phi(r) r^2 dr \quad (8)$$

It is easy to show that  $\int_0^R \phi(r) r^2 dr = A_2 \lambda_D (\lambda_D (1 - \exp(-R/\lambda_D)) - R \exp(-R/\lambda_D)) \rightarrow 0$  as  $R \rightarrow 0$  and  $\Phi \rightarrow 4\pi A_2$  as  $R \rightarrow 0$ . Thus, the Gauss' law ( $\Phi = Q_{\text{enc}}/\epsilon_0$ ) results in  $A_2 = Q/4\pi\epsilon_0$  and hence

$$\boxed{\phi(r) = \frac{Q}{4\pi\epsilon_0 r} \exp(-r/\lambda_D)}$$

c) The potential  $\phi_0(r)$  due to  $Q$  in the absence of plasma is trivially given by

$$\phi_0(r) = \frac{Q}{4\pi\epsilon_0 r} \quad (9)$$

and thus  $\phi(r) = \phi_0(r) \exp(-r/\lambda_D)$ . Thus, the effect of the charge  $Q$  is screened by the surrounding mobile plasma and becomes negligible over length scales of the order of  $\lambda_D$  which can be taken to be an estimate of the screening length.

## Problem 2

- a)  $|\psi(r)|^2$  is the probability of finding the electron per unit volume at a distance  $r$  from the origin. Since the electron must be found somewhere in all space,

$$\int_0^{\infty} |\psi(r)|^2 4\pi r^2 dr = 1 \quad (10)$$

Using  $\psi(r) = A \exp(-r/a)$ , we obtain:

$$\boxed{A = \frac{1}{\sqrt{\pi a^3}}} \quad (11)$$

- b) Consider a small volume  $\Delta V$  in space. The probability of finding the electron in that volume is  $|\psi|^2 dV$ . Thus the average charge  $\Delta Q$  in this volume is:

$$\Delta Q = -e|\psi(r)|^2 \Delta V \quad (12)$$

and hence the average charge density  $\rho(r)$

$$\rho(r) = \frac{\Delta Q}{\Delta V} = -e|\psi(r)|^2 = -\frac{e}{\pi a^3} \exp(-2r/a) \quad (13)$$

To calculate the field at  $r$  due to this charge distribution, consider a spherical gaussian surface of radius  $r$  centered at the origin and apply the gauss' law:

$$E(r)4\pi r^2 = \frac{1}{\epsilon_0} \int_0^r \frac{e}{\pi a^3} \exp(-2R/a) 4\pi R^2 dR \quad (14)$$

which, after some simplification, gives:

$$\boxed{E(r) = -\frac{e}{4\pi\epsilon_0 r^2} \left[ 1 - \exp(-2r/a) \left( 1 + \frac{2r}{a} + \frac{2r^2}{a^2} \right) \right]} \quad (15)$$

For small  $r$ , we can expand  $E(r)$  into a taylor series in  $r$  (or use  $\exp(-2r/a) \approx 1 - 2r/a + 2r^2/a^2 - 4r^3/3a^3$ ) to obtain:

$$\boxed{E(r) \approx -\frac{er}{3\pi\epsilon_0 a^3}} \quad (16)$$

- c) Consider an external field  $E_{\text{ex}}$ . This displaces the electron cloud from positive nucleus. We ignore the distortion in the shape of electron cloud. If the electron cloud displaced by distance  $d$  from nucleus due to the field. In equilibrium, the net force on nucleus is 0 and hence ( $E_{\text{ex}}$  and  $E$  refer only to the magnitudes of the external electric field and the electric field due to the electron cloud):

$$E(d) = E_{\text{ex}} \quad (17)$$

For small  $d$ , using Eq. 16

$$\boxed{p = ed = 3\pi\epsilon_0 a^3 E_{\text{ex}}} \quad (18)$$

where  $p$  is the dipole formed by the atom. Therefore the polarisability is given by  $\alpha = 3\pi\epsilon_0 a^3$ .